# May not sum to total due to rounding: the probability of rounding errors 

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#### Abstract

Many published datasets use rounded data, either because the measurement process has limited precision, or to improve presentation in the publication. This introduces an error of up to half the rounding precision, but when several pieces of rounded data are added together, the error in the sum can be larger. If the actual sum is also known, but presented as rounded to the same level, then this can be visible as the rounded components may not add up to the rounded sum; it can also be clear when the components are presented as rounded percentages but these do not add up $100 \%$. Where this happens, publishers often notate tables with the warning May not sum to total due to rounding or something similar.

This note explores some of the probabilities of rounding errors and visible rounding differences in sums, and provides tables of probabilities of different rounding errors. If many of the individual pieces of data are as small as the rounding precision or smaller, then this may also introduce bias into the sum of the rounding errors: three theoretical examples are considered, together with some actual data showing visible rounding differences.


## 1 Cumulative rounding errors

Initially we shall assume that the rounding error on individual components, the difference between the rounded and unrounded figure, is independently,
continuously and uniformly distributed between $-\frac{1}{2}$ and $\frac{1}{2}$ times the rounding precision, with zero probability of being at the extremes. We can then find the probability of distribution of the error of the sum by convoluting uniform distributions. Unfortunately this soon becomes complicated, though perfectly manageable using integer arithmetic with an unlimited precision computer. Writing $F_{n}(x)=\operatorname{Prob}\left(\sum_{i=1}^{n} X_{i} \leq x\right)$ for the cumulative distribution when $n$ independent rounding errors are combined, we get for small $n$

$$
\left.\begin{array}{c}
F_{1}(x)=\frac{2 x+1}{2} \quad \text { when } \quad-\frac{1}{2} \leq x \leq \frac{1}{2} . \\
F_{2}(x)= \begin{cases}\frac{x^{2}+2 x+1}{-\frac{x^{2}+2 x+1}{2}} & \text { when }-1 \leq x \leq 0\end{cases} \\
\text { when } \quad 0 \leq x \leq 1 .
\end{array}\right\} \begin{array}{ll}
F_{3}(x)= \begin{cases}\frac{8 x^{3}+36 x^{2}+54 x+27}{48} & \text { when }-\frac{3}{2} \leq x \leq-\frac{1}{2} \\
\frac{-4 x^{3}+9 x+6}{12} & \text { when }-\frac{1}{2} \leq x \leq \frac{1}{2} \\
\frac{8 x^{3}-36 x^{2}+54 x+21}{48} & \text { when } \quad \frac{1}{2} \leq x \leq \frac{3}{2} .\end{cases} \\
F_{4}(x)=\left\{\begin{array}{lll}
\frac{x^{4}+8 x^{3}+24 x^{2}+36 x+16}{24} & \text { when }-2 \leq x \leq-1 \\
\frac{-3 x^{4}-8 x^{3}+16 x+12}{3 x^{4}-8 x^{2}+16 x+12} & \text { when }-1 \leq x \leq 0 \\
\frac{-x^{4}+8 x^{3}-24 x^{2}+36 x+8}{24} & \text { when } & 0 \leq x \leq 1
\end{array}\right. \\
\text { when } 1 \leq x \leq 2 .
\end{array} .
$$

$$
F_{5}(x)= \begin{cases}\frac{32 x^{5}+400 x^{4}+2000 x^{3}+5000 x^{2}+6250 x+3125}{3840} & \text { when }-\frac{5}{2} \leq x \leq-\frac{3}{2} \\ \frac{-64 x^{5}-400 x^{4}-800 x^{3}-200 x^{2}+1100 x+955}{1920} & \text { when }-\frac{3}{2} \leq x \leq-\frac{1}{2} \\ \frac{48 x^{5}-200 x^{3}+575 x+480}{90} & \text { when }-\frac{1}{2} \leq x \leq \frac{1}{2} \\ \frac{-64 x^{5}+400 x^{4}-800 x^{3}+200 x^{2}+1100 x+965}{1920} & \text { when } \quad \frac{1}{2} \leq x \leq \frac{3}{2} \\ \frac{32 x^{5}-400 x^{4}+2000 x^{3}-5000 x^{2}+6250 x+715}{3840} & \text { when } \quad \frac{3}{2} \leq x \leq \frac{5}{2}\end{cases}
$$

Some of these can be factorised or otherwise simplified, but this becomes more difficult as $n$ increases. An alternative is to use the central limit theorem, noting that the initial uniform distribution has mean 0 and variance $\frac{1}{12}$, and so (assuming independence) the sum of $n$ such random variables has mean 0 and variance $\frac{n}{12}$; for large $n$ and $x$, this indeed provides reasonable approximations, though they are proportionately poor in the tails.

So $F_{n}(x)$ is the probability that with $n$ rounded numbers, the error in the sum due to rounding is less than or equal to $x$ times the rounding precision. Clearly

$$
\begin{gathered}
F_{n}(x)=0 \text { when } x \leq-\frac{n}{2} \\
F_{n}(x)=1 \text { when } x \geq \frac{n}{2}
\end{gathered}
$$

and the extreme events $\sum_{i=1}^{n} X_{i}= \pm \frac{n}{2}$ are possible but of zero probability.
In $F_{n}(x), x$ is signed; more often, we may be interested in whether the absolute value of the error in the sum is less than or equal to $x$ times the rounding precision for non-negative $x$. Calling this $G_{n}(x)$, we have

$$
G_{n}(x)=F_{n}(x)-F_{n}(-x)=2 F_{n}(x)-1 .
$$

and so $G_{n}(x)=1$ when $x \geq \frac{n}{2}$.
Rounding just one number gives a rounding error that is less than or equal to half the rounding precision. The probability this remains the case with the sum of $n$ rounded numbers is $G_{n}\left(\frac{1}{2}\right)$, while the probability that the error in the sum is less than the rounding precision is $G_{n}(1)$.

Snedecor and Cochran[1] provide an example problem along these lines:
Example 4.8.8. When measurements are rounded to the nearest whole number, it can often be assumed that the error due to rounding is equally likely to lie anywhere between -0.5 and +0.5 . That is, rounding errors follow a uniform distribution between the limits -0.5 and +0.5 . From theory, this distribution has $\mu=0$, $\sigma=1 / \sqrt{12}=0.29$. If 100 independent, rounded measurements are added, what is the probability that the error in the total due to rounding does not exceed 5 in absolute value? Ans. $P=0.916$.

Clearly the problem is asking us to find $G_{100}(5)$. Calculating the hundred convolutions gives the precise probability of

89120254737023412678295281505362351118315198761494175906 148647751101687378532185619182956244927133663945989014866250 791030484901069916977940250347876393 / 97214807754108492376 770040475277813011162466942064189029784337390851666659614495 425980691641829706548180935334195024748178317928038400000000 000000000000 , about 0.91674 .

The hint about the value of $\sigma$ and the position in the book suggest that problem does not expect such precision and instead expects use of the central limit theorem as a reasonable approximation. For a Gaussian distribution, the probability of being within an interval 2.9 standard deviations either side of the mean is about 0.91532 , while being more precise about the square root of $\frac{1}{12}$ would give a probability of about 0.91674 , suggesting that the provided answer faces some rounding issues of its own.

## 2 Visible rounding differences

That kind of analysis gives some indication of the probable sizes of errors introduced to a sum as a result of rounding. However, the actual error is not generally visible to external observers. But when individual items and the total are each rounded after the calculation of the unrounded sum, it is sometimes possible to observe that the rounded total is not equal to the sum of the rounded items. To remind readers of this and to avoid unnecessary enquires, many tables include the caveat May not sum to total due to rounding or something similar.

The analysis of the possible difference between sum of the rounded parts and the separately rounded total is similar to that before, but this time the rounded total must be an integer multiple of the rounding precision. So the probability that no error is visible is $F_{n}\left(\frac{1}{2}\right)-F_{n}\left(-\frac{1}{2}\right)=G_{n}\left(\frac{1}{2}\right)$. If the Snedecor and Cochran example had been reworded to ask What is the probability that the difference between the sum of the rounded individual measurements and the rounded total does not exceed 5 in absolute value? then the answer would be $G_{100}\left(5 \frac{1}{2}\right)$ or about 0.94325 . We could go further and look at the probabilities of particular values of the signed difference between the sum of $n$ rounded parts and the rounded total: if this difference was $d$ times the rounding precision and the probability was $H_{n}(d)$ then

$$
H_{n}(d)=F_{n}\left(d+\frac{1}{2}\right)-F_{n}\left(d-\frac{1}{2}\right) .
$$

## 3 Visible percentage differences

In many tables of sums, values themselves are not shown, and instead the parts are shown as percentages of the total at 100. Depending on how many
places are shown, this may involve rounding with more or less precision. Clearly the independence assumption used before has been lost, and the results will not be the same. To illustrate this, if there are only two parts, they can produce a visible rounding difference in the sum, but they cannot as percentages except in extreme cases: for example, 32.343 and 44.234 summing to 76.577 produces a difference when rounded but $42.236 \%$ and $57.764 \%$ summing to $100 \%$ does not. $42.5 \%$ and $57.5 \%$ might produce a visible rounding error depending on the rounding method, but this is by assumption a zero probability event.

Fortunately, it is possible to handle find a reasonable first approximation to using rounded percentages despite losing independence. If we only look at the fractional part of the rounding error for one part, our assumption is that this is uniform on $\left[-\frac{1}{2}, \frac{1}{2}\right]$ times the rounding precision, as shown by the derivative of $F_{1}(x)$. But for the fractional part it is also uniform on the same interval (modulo 1) for two parts, and by induction it is therefore also uniform on the same interval (modulo 1) for $n-1$ parts.

So if a further part is needed to reach the unrounded sum of $100 \%$, this further part will produce a fractional rounding error of the same magnitude but opposite in sign to fractional rounding error of the sum of the first $n-1$ rounded parts, and so we can take that it as having the same distribution as each of the others. Since this rounding error is less in magnitude than $\frac{1}{2}$ of the rounding precision, it does not add anything to the visible rounding difference caused by the first $n-1$ parts, and this means that if the signed difference between the sum of the $n$ rounded parts and $100 \%$ was $d$ times the rounding precision and the probability was $J_{n}(d)$ then

$$
J_{n}(d)=H_{n-1}(d) .
$$

## 4 Using the central limit theorem

Tables for $F_{n}(x), G_{n}(x)$ and $H_{n}(d)$, and so implicitly $J_{n}(d)$, are shown in later sections. But since these are based on the sum of independent identical distributions with finite variances, we can use the central limit theorem to use a Gaussian distribution to approximate these values.

The mean of one rounding error is 0 with a standard deviation of $\sqrt{\frac{1}{12}}$ times the rounding precision. So the sum of $n$ parts also has mean 0 and standard deviation $\sqrt{\frac{n}{12}}$. Writing $\Phi(x)$ for the cumulative distribution function
of a standard Gaussian random variable with mean 0 and standard deviation 1 , and $\phi(x)$ as the density, we then have

$$
F_{n}(x) \approx \Phi\left(\sqrt{\frac{12}{n}} x\right)
$$

and thus

$$
G_{n}(x) \approx 2 \Phi\left(\sqrt{\frac{12}{n}} x\right)-1
$$

and

$$
H_{n}(d) \approx \Phi\left(\sqrt{\frac{12}{n}}\left(d+\frac{1}{2}\right)\right)-\Phi\left(\sqrt{\frac{12}{n}}\left(d-\frac{1}{2}\right)\right) \approx \sqrt{\frac{12}{n}} \phi\left(\sqrt{\frac{12}{n}} d\right)
$$

Using this Gaussian approximation for $F_{n}(x)$ produces values which have a difference of less than 0.001 from the correct figure for $|x|>2$ or $n>11$, but requires much larger values to achieve better levels of accuracy. Since the approximations for $G_{n}(x)$ are twice as inaccurate, to be sure of figures which are within 0.001 of the correct figure, these become $x>3$ or $n>56$. The approximation for $G_{100}(5)$ is within $10^{-7}$ of the correct figure, but this is fortuitous: for $G_{100}(2)$ the approximation is more than $5 \times 10^{-4}$ too high, while for $G_{100}(7)$ the approximation is over $1.4 \times 10^{-4}$ too low.

As magnitudes, the Gaussian approximations are dreadful in the tails: for example with $|d| \geq 39$ and $n \leq 100$ they give magnitudes for $H_{n}(d)$ more than $10^{11}$ times the true figure, but this is in terms of very small probabilities as the true figure for $H_{100}(39)$ is about $1.244 \times 10^{-52}$ while the approximations suggest values of around $7 \times 10^{-41}$ or $3 \times 10^{-41}$.

The second Gaussian approximation given for $H_{n}(d)$ is notably weaker for $H_{n}(0)$, the probability of no visible rounding error. Better than either appears to be something like

$$
H_{n}(0)=G_{n}\left(\frac{1}{2}\right) \approx \sqrt{\frac{6}{\pi(n+1.3)}}
$$

and so

$$
J_{n}(0) \approx \sqrt{\frac{6}{\pi(n+0.3)}}
$$

## 5 Rounding a geometric series

It is easy to find cases of rounding large numbers of small parts. Take for example this geometric series

$$
1+\frac{9}{10}+\frac{81}{100}+\frac{729}{1000}+\frac{6561}{10000}+\cdots=10
$$

With a rounding precision of 1 , the rounded terms are just the sum of seven ones giving a sum of the rounded parts of 7 , and so a rounding difference of -3 . But taking a rounding precision of 0.01 , the first 51 parts each round to a positive number, and their sum is 9.98 , giving a rounding difference of -2 times the rounding precision. Going further, a rounding precision of 0.000001 requires the sum of 138 rounded parts making 10.000001 with a visible rounding difference of +1 times the rounding precision. So the rounding difference is variable; as it also depends on the factor in the geometric series, we may be able to approximate some kind of expected rounding difference.

If we have a series of the form

$$
\sum_{j=1}^{\infty} b k^{j}
$$

for some constants $b>0$ and $0<k<1$, and use a rounding precision of $p>0$ with $b$ substantially greater than $p$, then the terms which round to exactly $p$ lie in the interval $\left[\frac{1}{2} p, \frac{3}{2} p\right]$. We can expect about $\frac{\log \left(\frac{3 p}{2}\right)-\log \left(\frac{p}{2}\right)}{-\log (k)}=\frac{\log (3)}{-\log (k)}$ of them, and similarly about $\frac{\log (2 i+1)-\log (2 i-1)}{-\log (k)}$ terms rounding to $i p$, and so on all the way up to $b$; positive rounding errors are more likely than negative ones. We can use Stirling's formula to find that the expected total rounding error associated with terms rounding to a positive number is about

$$
\begin{gathered}
\sum_{i=1}^{\frac{b}{p}} \int_{i-\frac{1}{2}}^{\min \left(i+\frac{1}{2}, \frac{b}{p}\right)} \frac{i-x}{-x \log _{e}(k)} d x \approx \frac{\frac{b}{p} \log _{e}\left(\frac{b}{p}\right)-\left(\frac{b}{p}-\frac{1}{2}\right)-\sum_{i=1}^{\frac{b}{p}} \log _{e}\left(i-\frac{1}{2}\right)}{-\log _{e}(k)} \\
\left.=\frac{\frac{1}{2}+\log _{e}\left(\left(\frac{4}{e} \frac{b}{p}\right)^{\frac{b}{p}}\left(\frac{b}{p}\right)!\right.}{\left(\frac{2 b}{p}\right)!}\right) \\
-\log _{e}(k)
\end{gathered} \frac{1-\log _{e}(2)}{-2 \log _{e}(k)}
$$

times the rounding precision.

For the terms which round to zero, the largest term lies in the interval $\left[\frac{k}{2} p, \frac{1}{2} p\right]$ and so the negative of their sum is the rounding error and lies between $\frac{-k}{2(1-k)}$ and $\frac{-1}{2(1-k)}$ times the rounding precision. Using a similar assumption for the distribution to that implicitly assumed for the terms rounding to positive numbers (not that it makes much difference at this stage except for neatness), a reasonable central figure to take is $\frac{1}{2 \log _{e}(k)}$ times the rounding precision. This is negative and larger than the approximate expectation for the rounding error for terms rounding to positive numbers.

So the overall expectation for the rounding error for the sum of a geometric series is about

$$
\frac{\log (2)}{2 \log (k)}
$$

times the rounding precision, and is negative since $k<1$; given its relatively small size and the variability of rounding errors, many actual rounded geometric series can have a positive rounding error. For $k=\frac{9}{10}$ shown in the example at the start of this section, it suggests a figure of roughly -3.3 times the rounding precision. With tighter precision the expected rounding error tends to get smaller in absolute terms, but its expected value does not change substantially as a multiple of the rounding precision.

## 6 Rounding a power-law series

The previous example showed a case where the expected rounding error was small. It is possible to make it much larger: take for example this power-law series, the Basel problem calculating $\zeta(2)$ :

$$
1+\frac{1}{4}+\frac{1}{9}+\frac{1}{16}+\frac{1}{25}+\cdots=\frac{\pi^{2}}{6} \approx 1.644934
$$

With a rounding precision of 1 , the rounded terms are just $1+0+0+0+0+\cdots$ giving a sum of the rounded parts of 1 , while rounding the actual sum gives 2. But taking a rounding precision of 0.01 , the first fourteen parts round to a positive figure, and their sum is 1.59 while rounding the actual sum gives 1.64 giving a rounding difference of -5 times the rounding precision. Going further, a rounding precision of 0.000001 requires the sum of 1414 rounded parts to give 1.644332 giving a visible rounding difference of -602 times the rounding precision.

So making the rounding precision smaller tends to make the rounding error and the visible rounding difference a more negative multiple of the rounding precision. The exponent of the power law series also has an impact and here we used -2 : it is no surprise that the rounding error gets worse more quickly if the exponent is -1.5 than if it is -4 as we have more parts close to the rounding precision. If we have a series of the form

$$
\sum_{j=1}^{\infty} b j^{-k}
$$

for some constants $b>0$ and $k>1$, and use a rounding precision of $p>0$, then some experimentation suggests that the expected rounding error and visible rounding difference for a rounded power-law series might be something of the order of

$$
-0.6(k-1)^{-1.2}\left(\frac{b}{p}\right)^{\frac{1}{k}}
$$

times the rounding precision. It will not be that precise value both because the experimentation did not deliver a precise result and because of the natural distribution of errors, but for small $k$ and large $\frac{b}{p}$ it will dominate. Taking $b=1$ and $k=2$ as in the example at the beginning of this section, it gives expected rounding errors of -0.6 when the rounding precision is $1 ;-6$ times the rounding precision when that is 0.01 ; and -600 times the rounding precision when that is 0.000001 . With tighter precision the expected rounding error tends to get smaller in absolute terms, but tends to get bigger in magnitude as a multiple of the rounding precision.

## 7 A realistic distribution for percentages

The earlier section on percentages suggested it is possible to have identical locally uniform distributions for the parts of $100 \%$, that is to say of 1 . But at a global level, these is less realistic. It might be more plausible to consider a distribution which has a constant density subject to the constraints that each of the $n$ parts is non-negative and their sum is 1 . The density turns out to be $(n-1)$ !. It produces the following conditional distributions where $k$ parts are known: if we label the known parts in any order $A_{i}$ with values $a_{i}$ for $1 \leq i \leq k$, and the unknown parts in any order $A_{k+j}$ for $1 \leq j \leq n-k-1$,
we have

$$
\begin{gathered}
\operatorname{Prob}\left(A_{k+j} \leq x \mid A_{1}=a_{1}, \ldots, A_{k}=a_{k}\right)=1-\left(1-\frac{x}{1-\sum_{i=1}^{k} a_{i}}\right)^{n-k-1} \\
\quad \text { when } 0 \leq x \leq 1-\sum_{i=1}^{k} a_{i}
\end{gathered}
$$

and finally when $n-1$ parts are known

$$
A_{n}=1-\sum_{i=1}^{n-1} a_{i}
$$

so each part is identically distributed. Even if the overall density is constant, each part is more likely to be smaller than larger, as its marginal probability density is

$$
f(x)=(n-1)(1-x)^{n-2} \text { for } 0 \leq x \leq 1
$$

which is decreasing in $x$ so the signed rounding error as the difference between the rounded and the unrounded value of each part is therefore more likely to be negative than positive; the expected value of each part is of course $\frac{1}{n}$.

Imagine there are just three parts and the rounding precision is $50 \%$, so each part is rounded to $0 \%, 50 \%$ or $100 \%$. This is unrealistic, but is designed to demonstrate the idea simply enough to fit on the page while conveying the calculation. So a part will produce negative rounding error if its unrounded part is between $0 \%$ and $25 \%$ or between $50 \%$ or $75 \%$, and positive rounding error otherwise. But there may be offsetting rounding errors from the other parts. To get a negative visible rounding difference from the three rounded parts, two must round to $0 \%$ and one to $50 \%$ adding to $50 \%$, while to get a positive visible rounding difference each must round to $50 \%$ adding to $150 \%$; with more parts or a smaller rounding precision there would be many more possibilities. The probability of a negative visible rounding difference is
$\int_{0}^{\frac{1}{4}} \int_{\frac{1}{4}-x}^{\frac{1}{4}} 2 d y d x+\int_{0}^{\frac{1}{4}} \int_{\frac{3}{4}-x}^{\frac{3}{4}} 2 d y d x+\int_{\frac{1}{2}}^{\frac{3}{4}} \int_{\frac{3}{4}-x}^{\frac{1}{4}} 2 d y d x=\frac{1}{16}+\frac{1}{16}+\frac{1}{16}=\frac{3}{16}$
while the probability of a positive visible rounding difference is

$$
\int_{\frac{1}{4}}^{\frac{1}{2}} \int_{\frac{1}{4}}^{\frac{3}{4}-x} 2 d y d x=\frac{1}{16} .
$$

The earlier uniform approximation gave $J_{3}(-1)=J_{3}(1)=\frac{1}{8}$. With this more realistic distribution, we find that with rounded percentages there is a higher probability of the sum of the rounded parts being less than $100 \%$, as expected.

That example used a very large rounding precision. Using a more obvious rounding precision of $1 \%$ would give probabilities of negative and positive visible rounding differences of $\frac{101}{800}$ and $\frac{99}{800}$ respectively; using $0.1 \%$ would give $\frac{1001}{8000}$ and $\frac{999}{8000}$ respectively, with an obvious pattern. This justifies the earlier approximation as having been reasonable, but only when the rounding precision is small compared with the average size of the parts.

With a rounding precision of $p$, we can work out the expected rounding error for each part and multiply by $n$ to find the bias or expected total rounding error or visible rounding difference to be
$\frac{n}{p} \sum_{j=0}^{\frac{1}{p}} \int_{\max \left(0, j p-\frac{1}{2}\right)}^{\min \left(1, j p+\frac{1}{2}\right)}(j p-x)(n-1)(1-x)^{n-2} d x=-\frac{1}{p}+n\left(\frac{p}{2}\right)^{n-1} \sum_{j=1}^{\frac{1}{p}}(2 j-1)^{n-1}$
times the rounding precision. This expected rounding error will be negative: for small $n p$ with relatively few parts compared with the precision, this will be minor and close to $-\frac{n(n-1) p}{24}$ times the rounding precision; for large $n p$ it will tend towards $-\frac{1}{p}+n\left(1-\frac{p}{2}\right)^{n-1}$ times the rounding precision, as almost all the parts will round to 0 and so the sum of the rounded parts will be close to 0 . For example, with 100 parts and a rounding precision of $0.1 \%$ the expected rounding error is about -0.41 times the rounding precision so we expect the sum of the rounded parts to be close to $100.0 \%$, but then allowing for the dispersion of $J_{100}(d)$; with 1000 parts and a rounding precision of $1 \%$ the expected rounding error is about -93.3 times the rounding precision so we expect the sum of the rounded parts to be close to a hopeless $7 \%$, again with some dispersion. With a fixed number of parts, tighter precision tends to reduce the magnitude of expected rounding error as a multiple of the rounding precision.

## 8 An real example: reported vote shares

Two different methods have been put forward for estimating visible rounding differences for percentages by making assumptions about plausible distributions for rounding errors. In reality, distributions will not exactly follow
either of these assumptions, and this can affect the results though possibly not by enough to notice. This example illustrates such a case.

The Electoral Commission[2][3] published the results of the United Kingdom 2001 and 2005 parliamentary general elections showing the votes won by each candidate in each constituency. It also published the share of each candidate's vote in their constituency rounded to $0.1 \%$, and in some constituencies the sums of these showed visible percentage differences. For example, in Aberavon in 2001, the votes for the seven candidates were 19063, 2955, 2933, 2296, 1960, 727 and 256, making a total of 30190 so the vote shares were shown as $63.1 \%, 9.8 \%, 9.7 \%, 7.6 \%, 6.5 \%, 2.4 \%$ and $0.8 \%$, summing to $99.9 \%$ giving a visible difference of $-0.1 \%$.

In 2001 there were 3319 candidates in 659 constituencies, implying a mean number of about 5.04 candidates per constituency. The number in individual constituencies ranged from three to nine: 45 constituencies had three candidates, 198 had four, 208 had five, 134 had six, 49 had seven, 20 had eight, and 5 had nine. 1177 of the candidates lost their $£ 500$ deposits by receiving less than a $5 \%$ share of the valid votes in their constituency.

The election four years later in 2005 saw more candidates: there were 3554 candidates in 646 constituencies after some boundaries in Scotland had been redrawn, implying a mean number of about 5.50 candidates per constituency. The number in individual constituencies ranged from three to fifteen: 21 constituencies had three candidates, 136 had four, 215 had five, 128 had six, 92 had seven, 34 had eight, 17 had nine, 2 had ten, and 1 (Sedgefield) had an astonishing fifteen. 1385 of the candidates lost their deposits.

If we used the symmetric numbers for $J_{n}(d)$, weighted for the number of candidates per constituency, we could predict the number of constituencies with particular visible rounding differences. This table compares the actual distribution of sums of rounded vote percentages with these predictions.

| Sum of rounded <br> percentages | Actual <br> in 2001 | Predicted <br> for 2001 | Actual <br> in 2005 | Predicted <br> for 2005 |
| :---: | ---: | ---: | ---: | ---: |
| $99.7 \%$ | 0 | 0.01 | 0 | 0.03 |
| $99.8 \%$ | 3 | 3.06 | 8 | 4.58 |
| $99.9 \%$ | 144 | 125.97 | 117 | 129.48 |
| $100.0 \%$ | 398 | 400.93 | 376 | 377.82 |
| $100.1 \%$ | 111 | 125.97 | 138 | 129.48 |
| $100.2 \%$ | 3 | 3.06 | 7 | 4.58 |
| $100.3 \%$ | 0 | 0.01 | 0 | 0.03 |

$\chi^{2}$-tests with four degrees of freedom would not suggest the actual figures were significantly different from the predictions.

The distribution of vote shares of candidates was not consistent with the constant density percentage distribution assumed earlier: that would have predicted about 649.3 and 766.1 lost deposits in 2001 and 2005 respectively, just over half of the actual values which reflected the large number of fringe candidates. Despite that, the number of candidates who had such a low share of the vote that it rounded to $0.0 \%$ was smaller than would have been predicted: the constant density distribution would have predicted about 7.2 and 8.6 while the actual figures in the two elections were 0 and 4 . So the expected negative bias in the visible rounding differences might be overstated by the constant density percentage distribution, and it was in any case small since the predictions add up to about a net $-0.06 \%$ and $-0.07 \%$ across the all constituencies in each of the two elections; in fact the total net visible rounding differences were $-3.3 \%$ and $+1.9 \%$ in 2001 and 2005 respectively, with the natural dispersion of results overwhelming any bias.

Finally, the results also showed the majority in each constituency: the difference in votes between the most and second most popular candidates and similarly as a proportion of the total valid votes. As a vote share, this too showed visible rounding differences, and although it involves a subtraction rather than an addition, a similar analysis should be possible, leading to a distribution similar to $H_{2}(d)$ of $\frac{1}{8}, \frac{3}{4}$ and $\frac{1}{8}$; we use $H_{n}(d)$ rather than $J_{n}(d)$ since the two percentages are not constrained to come to a particular figure. Again taking Aberavon in 2001 as an example, the top two vote shares were given as $63.1 \%$ and $9.8 \%$ but the majority of 16108 votes was shown as $53.4 \%$, giving a visible difference of $-0.1 \%$. The following table compares the actual and predicted visible rounding differences in the majorities.

| Visible <br> differences <br> in majorities | Actual <br> in 2001 | Predicted <br> for 2001 | Actual <br> in 2005 | Predicted <br> for 2005 |
| :---: | ---: | :---: | ---: | :---: |
| $-0.1 \%$ | 90 | 82.375 | 93 | 80.75 |
| $0.0 \%$ | 480 | 494.25 | 478 | 484.5 |
| $+0.1 \%$ | 89 | 82.375 | 75 | 80.75 |

Again $\chi^{2}$-tests, this time with two degrees of freedom, would not suggest the actual figures were significantly different from the predictions. If a Gaussian approximation had been used instead, they would seem significant.

## 9 Table of $F_{n}(x)$

The cumulative distribution $F_{n}(x)$ : the probability that the signed rounding error of the sum of $n$ rounded terms is less than or equal to $x$ times the rounding precision. By symmetry $F_{n}(0)=\frac{1}{2}$ and $F_{n}(-x)=1-F_{n}(x)$.

| $n$ | $F_{n}(0.5)$ | $F_{n}(1)$ | $F_{n}(1.5)$ | $F_{n}(2)$ | $F_{n}(2.5)$ | $F_{n}(3)$ | $F_{n}(3.5)$ | $F_{n}(4)$ | $F_{n}(4.5)$ | $F_{n}(5)$ | $F_{n}(5.5)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1 |  |  |  |  |  |  |  |  |  |  |
| 2 | 0.875 | 1 |  |  |  |  |  |  |  |  |  |
| 3 | 0.83333 | 0.97917 | 1 |  |  |  |  |  |  |  |  |
| 4 | 0.79948 | 0.95833 | 0.99740 | 1 |  |  |  |  |  |  |  |
| 5 | 0.775 | 0.93802 | 0.99167 | 0.99974 | 1 |  |  |  |  |  |  |
| 6 | 0.75551 | 0.91944 | 0.98431 | 0.99861 | 0.99998 | 1 |  |  |  |  |  |
| 7 | 0.73968 | 0.90260 | 0.97599 | 0.99662 | 0.99980 | 1.00000 | 1 |  |  |  |  |
| 8 | 0.72646 | 0.88738 | 0.96724 | 0.99385 | 0.99937 | 0.99998 | 1.00000 | 1 |  |  |  |
| 9 | 0.71521 | 0.87360 | 0.95836 | 0.99044 | 0.99861 | 0.99989 | 1.00000 | 1.00000 | 1 |  |  |
| 10 | 0.70548 | 0.86110 | 0.94955 | 0.98654 | 0.99753 | 0.99972 | 0.99998 | 1.00000 | 1.00000 | 1 |  |
| 11 | 0.69696 | 0.84970 | 0.94092 | 0.98226 | 0.99613 | 0.99943 | 0.99995 | 1.00000 | 1.00000 | 1.00000 | 1 |
| 12 | 0.68942 | 0.83927 | 0.93255 | 0.97772 | 0.99442 | 0.99899 | 0.99988 | 0.99999 | 1.00000 | 1.00000 | 1.00000 |
| 13 | 0.68269 | 0.82969 | 0.92447 | 0.97300 | 0.99245 | 0.99841 | 0.99976 | 0.99998 | 1.00000 | 1.00000 | 1.00000 |
| 14 | 0.67662 | 0.82085 | 0.91670 | 0.96817 | 0.99024 | 0.99767 | 0.99958 | 0.99995 | 1.00000 | 1.00000 | 1.00000 |
| 15 | 0.67112 | 0.81267 | 0.90924 | 0.96328 | 0.98784 | 0.99678 | 0.99934 | 0.99990 | 0.99999 | 1.00000 | 1.00000 |
| 16 | 0.66610 | 0.80506 | 0.90209 | 0.95837 | 0.98527 | 0.99575 | 0.99902 | 0.99983 | 0.99998 | 1.00000 | 1.00000 |
| 17 | 0.66150 | 0.79798 | 0.89524 | 0.95346 | 0.98255 | 0.99457 | 0.99863 | 0.99972 | 0.99996 | 1.00000 | 1.00000 |
| 18 | 0.65727 | 0.79136 | 0.88868 | 0.94860 | 0.97973 | 0.99327 | 0.99815 | 0.99959 | 0.99993 | 0.99999 | 1.00000 |
| 19 | 0.65335 | 0.78515 | 0.88239 | 0.94379 | 0.97681 | 0.99185 | 0.99759 | 0.99941 | 0.99988 | 0.99998 | 1.00000 |
| 20 | 0.64971 | 0.77932 | 0.87637 | 0.93904 | 0.97383 | 0.99032 | 0.99695 | 0.99920 | 0.99983 | 0.99997 | 1.00000 |
| 21 | 0.64631 | 0.77383 | 0.87059 | 0.93438 | 0.97079 | 0.98869 | 0.99624 | 0.99894 | 0.99975 | 0.99995 | 0.99999 |
| 22 | 0.64314 | 0.76865 | 0.86505 | 0.92980 | 0.96771 | 0.98698 | 0.99545 | 0.99863 | 0.99965 | 0.99993 | 0.99999 |
| 23 | 0.64016 | 0.76374 | 0.85973 | 0.92531 | 0.96461 | 0.98519 | 0.99458 | 0.99828 | 0.99953 | 0.99989 | 0.99998 |
| 24 | 0.63737 | 0.75910 | 0.85462 | 0.92091 | 0.96149 | 0.98334 | 0.99365 | 0.99788 | 0.99939 | 0.99985 | 0.99997 |
| 25 | 0.63473 | 0.75468 | 0.84971 | 0.91660 | 0.95837 | 0.98143 | 0.99265 | 0.99744 | 0.99922 | 0.99979 | 0.99995 |
| 26 | 0.63224 | 0.75049 | 0.84499 | 0.91240 | 0.95525 | 0.97946 | 0.99159 | 0.99694 | 0.99902 | 0.99973 | 0.99993 |
| 27 | 0.62988 | 0.74649 | 0.84044 | 0.90828 | 0.95214 | 0.97745 | 0.99047 | 0.99641 | 0.99880 | 0.99965 | 0.99991 |
| 28 | 0.62765 | 0.74268 | 0.83606 | 0.90427 | 0.94905 | 0.97541 | 0.98930 | 0.99582 | 0.99855 | 0.99955 | 0.99988 |
| 29 | 0.62552 | 0.73904 | 0.83183 | 0.90034 | 0.94597 | 0.97334 | 0.98808 | 0.99520 | 0.99827 | 0.99944 | 0.99984 |
| 30 | 0.62350 | 0.73555 | 0.82776 | 0.89651 | 0.94292 | 0.97124 | 0.98682 | 0.99453 | 0.99796 | 0.99932 | 0.99980 |
| 31 | 0.62158 | 0.73222 | 0.82382 | 0.89277 | 0.93990 | 0.96912 | 0.98552 | 0.99383 | 0.99762 | 0.99917 | 0.99974 |
| 32 | 0.61974 | 0.72902 | 0.82002 | 0.88912 | 0.93691 | 0.96698 | 0.98417 | 0.99308 | 0.99725 | 0.99901 | 0.99968 |
| 33 | 0.61798 | 0.72595 | 0.81634 | 0.88555 | 0.93395 | 0.96484 | 0.98280 | 0.99230 | 0.99686 | 0.99883 | 0.99961 |
| 34 | 0.61630 | 0.72300 | 0.81279 | 0.88206 | 0.93102 | 0.96269 | 0.98140 | 0.99148 | 0.99643 | 0.99864 | 0.99953 |
| 35 | 0.61468 | 0.72016 | 0.80935 | 0.87866 | 0.92813 | 0.96053 | 0.97996 | 0.99063 | 0.99598 | 0.99842 | 0.99944 |
| 36 | 0.61314 | 0.71742 | 0.80601 | 0.87534 | 0.92528 | 0.95837 | 0.97851 | 0.98975 | 0.99550 | 0.99819 | 0.99933 |
| 37 | 0.61165 | 0.71479 | 0.80278 | 0.87209 | 0.92246 | 0.95621 | 0.97703 | 0.98884 | 0.99500 | 0.99793 | 0.99922 |
| 38 | 0.61022 | 0.71225 | 0.79965 | 0.86892 | 0.91969 | 0.95405 | 0.97553 | 0.98790 | 0.99446 | 0.99766 | 0.99909 |
| 39 | 0.60884 | 0.70979 | 0.79661 | 0.86582 | 0.91695 | 0.95190 | 0.97402 | 0.98694 | 0.99391 | 0.99737 | 0.99895 |
| 40 | 0.60752 | 0.70742 | 0.79366 | 0.86279 | 0.91424 | 0.94976 | 0.97249 | 0.98595 | 0.99333 | 0.99706 | 0.99880 |
| 41 | 0.60624 | 0.70513 | 0.79079 | 0.85983 | 0.91158 | 0.94763 | 0.97095 | 0.98494 | 0.99273 | 0.99673 | 0.99864 |
| 42 | 0.60501 | 0.70291 | 0.78801 | 0.85694 | 0.90896 | 0.94550 | 0.96939 | 0.98391 | 0.99210 | 0.99639 | 0.99846 |
| 43 | 0.60381 | 0.70076 | 0.78530 | 0.85411 | 0.90637 | 0.94339 | 0.96783 | 0.98286 | 0.99145 | 0.99602 | 0.99828 |
| 44 | 0.60266 | 0.69868 | 0.78266 | 0.85134 | 0.90382 | 0.94129 | 0.96627 | 0.98179 | 0.99079 | 0.99564 | 0.99807 |
| 45 | 0.60155 | 0.69666 | 0.78010 | 0.84863 | 0.90131 | 0.93921 | 0.96469 | 0.98071 | 0.99010 | 0.99524 | 0.99786 |
| 46 | 0.60047 | 0.69470 | 0.77760 | 0.84597 | 0.89884 | 0.93714 | 0.96311 | 0.97961 | 0.98939 | 0.99482 | 0.99763 |
| 47 | 0.59942 | 0.69280 | 0.77517 | 0.84338 | 0.89640 | 0.93508 | 0.96153 | 0.97849 | 0.98867 | 0.99439 | 0.99739 |
| 48 | 0.59841 | 0.69096 | 0.77280 | 0.84084 | 0.89401 | 0.93304 | 0.95995 | 0.97736 | 0.98793 | 0.99394 | 0.99714 |
| 49 | 0.59743 | 0.68916 | 0.77049 | 0.83835 | 0.89164 | 0.93102 | 0.95837 | 0.97623 | 0.98717 | 0.99348 | 0.99688 |
| 50 | 0.59647 | 0.68742 | 0.76823 | 0.83591 | 0.88932 | 0.92901 | 0.95678 | 0.97508 | 0.98640 | 0.99300 | 0.99660 |

$F_{n}(x)$ continued: the probability that the signed rounding error of the sum of $n$ rounded terms is less than or equal to $x$ times the rounding precision.
$F_{n}(0)=\frac{1}{2}$ and $F_{n}(-x)=1-F_{n}(x)$.

| $n$ | $F_{n}(0.5)$ | $F_{n}(1)$ | $F_{n}(1.5)$ | $F_{n}(2)$ | $F_{n}(2.5)$ | $F_{n}(3)$ | $F_{n}$ | $F_{n}(4)$ | $F_{n}(4.5)$ | $F_{n}(5)$ | $F_{n}(5.5)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 51 | 0.59555 | 0.68572 | 0.76603 | 0.83352 | 0.88702 | 0.92702 | 0.95520 | 0.97392 | 0.98562 | 0.99250 | 0.99631 |
| 52 | 0.59465 | 0.68407 | 0.76388 | 0.83118 | 0.88477 | 0.92505 | 0.95362 | 0.97275 | 0.98482 | 0.99199 | 0.99601 |
| 53 | 0.59377 | 0.68246 | 0.76179 | 0.82889 | 0.88254 | 0.92309 | 0.95205 | 0.97157 | 0.98400 | 0.99147 | 0.99569 |
| 54 | 0.59292 | 0.68089 | 0.75974 | 0.82664 | 0.88035 | 0.92116 | 0.95048 | 0.97039 | 0.98318 | 0.99093 | 0.99536 |
| 55 | 0.59209 | 0.67936 | 0.75774 | 0.82443 | 0.87819 | 0.91924 | 0.94891 | 0.96921 | 0.98234 | 0.99038 | 0.99503 |
| 56 | 0.59128 | 0.67787 | 0.75578 | 0.82227 | 0.87607 | 0.91734 | 0.94734 | 0.96801 | 0.98149 | 0.98982 | 0.99468 |
| 57 | 0.59050 | 0.67642 | 0.75387 | 0.82015 | 0.87397 | 0.91546 | 0.94579 | 0.96682 | 0.98064 | 0.98924 | 0.99432 |
| 58 | 0.58973 | 0.67500 | 0.75200 | 0.81806 | 0.87191 | 0.91359 | 0.94424 | 0.96561 | 0.97977 | 0.98866 | 0.99394 |
| 59 | 0.58898 | 0.67361 | 0.75017 | 0.81602 | 0.86988 | 0.91175 | 0.94269 | 0.96441 | 0.97889 | 0.98806 | 0.99356 |
| 60 | 0.58825 | 0.67226 | 0.74838 | 0.81401 | 0.86787 | 0.90992 | 0.94115 | 0.96321 | 0.97801 | 0.98745 | 0.99317 |
| 61 | 0.58754 | 0.67094 | 0.74662 | 0.81204 | 0.86590 | 0.90812 | 0.93962 | 0.96200 | 0.97712 | 0.98683 | 0.99277 |
| 62 | 0.58685 | 0.66965 | 0.74491 | 0.81011 | 0.86395 | 0.90633 | 0.93810 | 0.96079 | 0.97622 | 0.98620 | 0.99235 |
| 63 | 0.58617 | 0.66839 | 0.74322 | 0.80821 | 0.86204 | 0.90456 | 0.93658 | 0.95958 | 0.97531 | 0.98557 | 0.99193 |
| 64 | 0.58551 | 0.66715 | 0.74157 | 0.80634 | 0.86015 | 0.90280 | 0.93508 | 0.95837 | 0.97440 | 0.98492 | 0.99150 |
| 65 | 0.58486 | 0.66594 | 0.73996 | 0.80451 | 0.85828 | 0.90107 | 0.93358 | 0.95716 | 0.97348 | 0.98426 | 0.99106 |
| 66 | 0.58423 | 0.66476 | 0.73838 | 0.80271 | 0.85645 | 0.89935 | 0.93209 | 0.95595 | 0.97256 | 0.98360 | 0.99061 |
| 67 | 0.58361 | 0.66360 | 0.73682 | 0.80094 | 0.85464 | 0.89765 | 0.93061 | 0.95474 | 0.97163 | 0.98293 | 0.99015 |
| 68 | 0.58300 | 0.66247 | 0.73530 | 0.79919 | 0.85285 | 0.89597 | 0.92914 | 0.95353 | 0.97070 | 0.98225 | 0.98968 |
| 69 | 0.58241 | 0.66136 | 0.73380 | 0.79748 | 0.85109 | 0.89431 | 0.92767 | 0.95233 | 0.96976 | 0.98157 | 0.98920 |
| 70 | 0.58183 | 0.66027 | 0.73234 | 0.79580 | 0.84935 | 0.89266 | 0.92622 | 0.95112 | 0.96883 | 0.98087 | 0.98872 |
| 71 | 0.58126 | 0.65920 | 0.73090 | 0.79414 | 0.84764 | 0.89103 | 0.92478 | 0.94992 | 0.96788 | 0.98017 | 0.98823 |
| 72 | 0.58071 | 0.65816 | 0.72948 | 0.79251 | 0.84595 | 0.88942 | 0.92334 | 0.94872 | 0.96694 | 0.97947 | 0.98773 |
| 73 | 0.58016 | 0.65713 | 0.72810 | 0.79091 | 0.84429 | 0.88783 | 0.92192 | 0.94753 | 0.96599 | 0.97876 | 0.98723 |
| 74 | 0.57963 | 0.65613 | 0.72673 | 0.78933 | 0.84264 | 0.88625 | 0.92050 | 0.94634 | 0.96504 | 0.97804 | 0.98671 |
| 75 | 0.57911 | 0.65514 | 0.72539 | 0.78778 | 0.84102 | 0.88469 | 0.91910 | 0.94515 | 0.96409 | 0.97732 | 0.98619 |
| 76 | 0.57859 | 0.65417 | 0.72408 | 0.78625 | 0.83942 | 0.88314 | 0.91770 | 0.94396 | 0.96314 | 0.97660 | 0.98567 |
| 77 | 0.57809 | 0.65322 | 0.72279 | 0.78475 | 0.83784 | 0.88161 | 0.91632 | 0.94278 | 0.96219 | 0.97587 | 0.98514 |
| 78 | 0.57759 | 0.65229 | 0.72151 | 0.78326 | 0.83628 | 0.88010 | 0.91494 | 0.94160 | 0.96123 | 0.97513 | 0.98460 |
| 79 | 0.57711 | 0.65138 | 0.72026 | 0.78180 | 0.83474 | 0.87860 | 0.91358 | 0.94043 | 0.96028 | 0.97439 | 0.98405 |
| 80 | 0.57663 | 0.65048 | 0.71904 | 0.78037 | 0.83323 | 0.87712 | 0.91222 | 0.93926 | 0.95932 | 0.97365 | 0.98351 |
| 81 | 0.57617 | 0.64959 | 0.71783 | 0.77895 | 0.83173 | 0.87565 | 0.91087 | 0.93810 | 0.95837 | 0.97291 | 0.98295 |
| 82 | 0.57571 | 0.64872 | 0.71664 | 0.77756 | 0.83025 | 0.87420 | 0.90954 | 0.93694 | 0.95741 | 0.97216 | 0.98239 |
| 83 | 0.57526 | 0.64787 | 0.71547 | 0.77618 | 0.82879 | 0.87276 | 0.90821 | 0.93578 | 0.95646 | 0.97141 | 0.98183 |
| 84 | 0.57482 | 0.64703 | 0.71432 | 0.77483 | 0.82734 | 0.87134 | 0.90690 | 0.93463 | 0.95550 | 0.97065 | 0.98126 |
| 85 | 0.57438 | 0.64621 | 0.71318 | 0.77349 | 0.82592 | 0.86993 | 0.90559 | 0.93349 | 0.95455 | 0.96990 | 0.98068 |
| 86 | 0.57395 | 0.64540 | 0.71207 | 0.77217 | 0.82451 | 0.86853 | 0.90429 | 0.93235 | 0.95360 | 0.96914 | 0.98011 |
| 87 | 0.57353 | 0.64460 | 0.71097 | 0.77088 | 0.82312 | 0.86715 | 0.90301 | 0.93121 | 0.95264 | 0.96838 | 0.97953 |
| 88 | 0.57312 | 0.64381 | 0.70989 | 0.76960 | 0.82175 | 0.86579 | 0.90173 | 0.93008 | 0.95169 | 0.96761 | 0.97894 |
| 89 | 0.57272 | 0.64304 | 0.70883 | 0.76833 | 0.82039 | 0.86444 | 0.90046 | 0.92896 | 0.95075 | 0.96685 | 0.97835 |
| 90 | 0.57232 | 0.64228 | 0.70778 | 0.76709 | 0.81905 | 0.86310 | 0.89921 | 0.92784 | 0.94980 | 0.96608 | 0.97776 |
| 91 | 0.57192 | 0.64153 | 0.70674 | 0.76586 | 0.81773 | 0.86177 | 0.89796 | 0.92673 | 0.94885 | 0.96531 | 0.97716 |
| 92 | 0.57154 | 0.64080 | 0.70572 | 0.76465 | 0.81642 | 0.86046 | 0.89672 | 0.92562 | 0.94791 | 0.96454 | 0.97656 |
| 93 | 0.57116 | 0.64007 | 0.70472 | 0.76345 | 0.81513 | 0.85916 | 0.89549 | 0.92452 | 0.94696 | 0.96377 | 0.97596 |
| 94 | 0.57078 | 0.63936 | 0.70373 | 0.76228 | 0.81385 | 0.85788 | 0.89427 | 0.92342 | 0.94602 | 0.96300 | 0.97535 |
| 95 | 0.57041 | 0.63866 | 0.70276 | 0.76111 | 0.81259 | 0.85660 | 0.89306 | 0.92233 | 0.94509 | 0.96223 | 0.97474 |
| 96 | 0.57005 | 0.63796 | 0.70180 | 0.75996 | 0.81134 | 0.85534 | 0.89186 | 0.92124 | 0.94415 | 0.96146 | 0.97413 |
| 97 | 0.56969 | 0.63728 | 0.70085 | 0.75883 | 0.81011 | 0.85409 | 0.89067 | 0.92016 | 0.94322 | 0.96069 | 0.97352 |
| 98 | 0.56934 | 0.63661 | 0.69992 | 0.75771 | 0.80889 | 0.85286 | 0.88949 | 0.91909 | 0.94229 | 0.95991 | 0.97290 |
| 99 | 0.56900 | 0.63595 | 0.69900 | 0.75661 | 0.80769 | 0.85163 | 0.88831 | 0.91802 | 0.94136 | 0.95914 | 0.97229 |
| 100 | 0.56865 | 0.63530 | 0.69809 | 0.75551 | 0.80649 | 0.85042 | 0.88715 | 0.91696 | 0.94043 | 0.95837 | 0.97167 |

## 10 Table of $G_{n}(x)$

$G_{n}(x)$ : the probability that the absolute rounding error of the sum of $n$ rounded terms is less than or equal to $x$ times the rounding precision. We have $G_{n}(0)=0$ since $G_{n}(x)=2 F_{n}(x)-1$.

| $n$ | $G_{n}(0.5)$ | $G_{n}(1)$ | $G_{n}(1.5)$ | $G_{n}(2)$ | $G_{n}(2.5)$ | $G_{n}(3)$ | $G_{n}(3.5)$ | $G_{n}(4)$ | $G_{n}(4.5)$ | $G_{n}(5)$ | $G_{n}(5.5)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1 |  |  |  |  |  |  |  |  |  |  |
| 2 | 0.75 | 1 |  |  |  |  |  |  |  |  |  |
| 3 | 0.66667 | 0.95833 | 1 |  |  |  |  |  |  |  |  |
| 4 | 0.59896 | 0.91667 | 0.99479 | 1 |  |  |  |  |  |  |  |
| 5 | 0.55 | 0.87604 | 0.98333 | 0.99948 | 1 |  |  |  |  |  |  |
| 6 | 0.51102 | 0.83889 | 0.96862 | 0.99722 | 0.99996 | 1 |  |  |  |  |  |
| 7 | 0.47937 | 0.80519 | 0.95198 | 0.99324 | 0.99960 | 1.00000 | 1 |  |  |  |  |
| 8 | 0.45292 | 0.77475 | 0.93448 | 0.98770 | 0.99873 | 0.99995 | 1.00000 | 1 |  |  |  |
| 9 | 0.43042 | 0.74720 | 0.91672 | 0.98088 | 0.99723 | 0.99979 | 0.99999 | 1.00000 | 1 |  |  |
| 10 | 0.41096 | 0.72220 | 0.89909 | 0.97307 | 0.99506 | 0.99944 | 0.99997 | 1.00000 | 1.00000 | 1 |  |
| 11 | 0.39393 | 0.69940 | 0.88185 | 0.96453 | 0.99225 | 0.99885 | 0.99990 | 1.00000 | 1.00000 | 1.00000 | 1 |
| 12 | 0.37884 | 0.67855 | 0.86511 | 0.95545 | 0.98884 | 0.99799 | 0.99976 | 0.99998 | 1.00000 | 1.00000 | 1.00000 |
| 13 | 0.36537 | 0.65938 | 0.84895 | 0.94601 | 0.98490 | 0.99682 | 0.99952 | 0.99995 | 1.00000 | 1.00000 | 1.00000 |
| 14 | 0.35324 | 0.64170 | 0.83341 | 0.93634 | 0.98049 | 0.99534 | 0.99917 | 0.99990 | 0.99999 | 1.00000 | 1.00000 |
| 15 | 0.34224 | 0.62533 | 0.81849 | 0.92656 | 0.97568 | 0.99356 | 0.99868 | 0.99980 | 0.99998 | 1.00000 | 1.00000 |
| 16 | 0.33221 | 0.61013 | 0.80418 | 0.91673 | 0.97053 | 0.99149 | 0.99805 | 0.99965 | 0.99996 | 1.00000 | 1.00000 |
| 17 | 0.32301 | 0.59596 | 0.79048 | 0.90693 | 0.96511 | 0.98914 | 0.99725 | 0.99945 | 0.99992 | 0.99999 | 1.00000 |
| 18 | 0.31453 | 0.58272 | 0.77736 | 0.89720 | 0.95945 | 0.98654 | 0.99630 | 0.99918 | 0.99986 | 0.99998 | 1.00000 |
| 19 | 0.30669 | 0.57031 | 0.76478 | 0.88757 | 0.95362 | 0.98370 | 0.99519 | 0.99883 | 0.99977 | 0.99996 | 1.00000 |
| 20 | 0.29941 | 0.55864 | 0.75274 | 0.87809 | 0.94765 | 0.98064 | 0.99391 | 0.99839 | 0.99965 | 0.99994 | 0.99999 |
| 21 | 0.29262 | 0.54766 | 0.74118 | 0.86876 | 0.94157 | 0.97739 | 0.99248 | 0.99787 | 0.99950 | 0.99990 | 0.99998 |
| 22 | 0.28628 | 0.53729 | 0.73010 | 0.85959 | 0.93542 | 0.97396 | 0.99089 | 0.99726 | 0.99930 | 0.99985 | 0.99997 |
| 23 | 0.28033 | 0.52749 | 0.71946 | 0.85061 | 0.92921 | 0.97039 | 0.98916 | 0.99656 | 0.99906 | 0.99978 | 0.99996 |
| 24 | 0.27473 | 0.51819 | 0.70924 | 0.84182 | 0.92298 | 0.96668 | 0.98729 | 0.99576 | 0.99878 | 0.99970 | 0.99994 |
| 25 | 0.26946 | 0.50937 | 0.69942 | 0.83321 | 0.91673 | 0.96285 | 0.98529 | 0.99487 | 0.99844 | 0.99959 | 0.99991 |
| 26 | 0.26448 | 0.50098 | 0.68997 | 0.82479 | 0.91050 | 0.95892 | 0.98317 | 0.99389 | 0.99805 | 0.99946 | 0.99987 |
| 27 | 0.25977 | 0.49298 | 0.68088 | 0.81657 | 0.90428 | 0.95491 | 0.98094 | 0.99281 | 0.99760 | 0.99930 | 0.99982 |
| 28 | 0.25530 | 0.48536 | 0.67211 | 0.80853 | 0.89809 | 0.95082 | 0.97860 | 0.99165 | 0.99710 | 0.99911 | 0.99976 |
| 29 | 0.25105 | 0.47808 | 0.66366 | 0.80069 | 0.89194 | 0.94667 | 0.97616 | 0.99040 | 0.99653 | 0.99889 | 0.99968 |
| 30 | 0.24701 | 0.47111 | 0.65551 | 0.79302 | 0.88585 | 0.94248 | 0.97364 | 0.98907 | 0.99592 | 0.99863 | 0.99959 |
| 31 | 0.24315 | 0.46444 | 0.64764 | 0.78554 | 0.87980 | 0.93824 | 0.97103 | 0.98765 | 0.99524 | 0.99835 | 0.99949 |
| 32 | 0.23948 | 0.45804 | 0.64004 | 0.77823 | 0.87382 | 0.93397 | 0.96835 | 0.98616 | 0.99450 | 0.99803 | 0.99936 |
| 33 | 0.23596 | 0.45190 | 0.63269 | 0.77110 | 0.86790 | 0.92968 | 0.96560 | 0.98460 | 0.99371 | 0.99767 | 0.99922 |
| 34 | 0.23259 | 0.44599 | 0.62558 | 0.76413 | 0.86205 | 0.92537 | 0.96279 | 0.98296 | 0.99286 | 0.99728 | 0.99906 |
| 35 | 0.22937 | 0.44031 | 0.61870 | 0.75733 | 0.85627 | 0.92105 | 0.95993 | 0.98126 | 0.99196 | 0.99684 | 0.99887 |
| 36 | 0.22627 | 0.43485 | 0.61203 | 0.75068 | 0.85056 | 0.91673 | 0.95701 | 0.97950 | 0.99100 | 0.99638 | 0.99867 |
| 37 | 0.22330 | 0.42958 | 0.60557 | 0.74419 | 0.84493 | 0.91242 | 0.95406 | 0.97768 | 0.98999 | 0.99587 | 0.99844 |
| 38 | 0.22044 | 0.42449 | 0.59930 | 0.73785 | 0.83937 | 0.90811 | 0.95106 | 0.97581 | 0.98893 | 0.99532 | 0.99818 |
| 39 | 0.21769 | 0.41959 | 0.59322 | 0.73165 | 0.83389 | 0.90381 | 0.94803 | 0.97388 | 0.98782 | 0.99474 | 0.99791 |
| 40 | 0.21504 | 0.41484 | 0.58732 | 0.72559 | 0.82849 | 0.89952 | 0.94497 | 0.97190 | 0.98666 | 0.99412 | 0.99761 |
| 41 | 0.21248 | 0.41026 | 0.58158 | 0.71967 | 0.82316 | 0.89525 | 0.94189 | 0.96988 | 0.98545 | 0.99347 | 0.99728 |
| 42 | 0.21001 | 0.40582 | 0.57601 | 0.71388 | 0.81791 | 0.89101 | 0.93879 | 0.96782 | 0.98420 | 0.99277 | 0.99693 |
| 43 | 0.20763 | 0.40152 | 0.57059 | 0.70821 | 0.81274 | 0.88678 | 0.93567 | 0.96572 | 0.98291 | 0.99204 | 0.99655 |
| 44 | 0.20533 | 0.39736 | 0.56532 | 0.70267 | 0.80764 | 0.88258 | 0.93253 | 0.96358 | 0.98157 | 0.99128 | 0.99615 |
| 45 | 0.20310 | 0.39332 | 0.56019 | 0.69725 | 0.80262 | 0.87841 | 0.92938 | 0.96141 | 0.98020 | 0.99048 | 0.99572 |
| 46 | 0.20094 | 0.38941 | 0.55520 | 0.69195 | 0.79768 | 0.87427 | 0.92623 | 0.95921 | 0.97879 | 0.98965 | 0.99527 |
| 47 | 0.19885 | 0.38561 | 0.55033 | 0.68676 | 0.79281 | 0.87016 | 0.92307 | 0.95698 | 0.97734 | 0.98878 | 0.99479 |
| 48 | 0.19682 | 0.38191 | 0.54559 | 0.68167 | 0.78801 | 0.86608 | 0.91990 | 0.95473 | 0.97586 | 0.98788 | 0.99428 |
| 49 | 0.19485 | 0.37832 | 0.54097 | 0.67670 | 0.78328 | 0.86203 | 0.91674 | 0.95245 | 0.97435 | 0.98695 | 0.99375 |
| 50 | 0.19295 | 0.37483 | 0.53646 | 0.67182 | 0.77863 | 0.85802 | 0.91357 | 0.95015 | 0.97281 | 0.98599 | 0.99320 |

$G_{n}(x)$ continued: the probability that the absolute rounding error of the sum of $n$ rounded terms is less than or equal to $x$ times the rounding precision. $G_{n}(0)=0$.

| $n$ | $G_{n}(0.5)$ | $G_{n}(1)$ | $G_{n}(1.5)$ | $G_{n}(2)$ | $G_{n}(2.5)$ | $G_{n}(3)$ | $G_{n}(3.5)$ | $G_{n}(4)$ | $G_{n}(4.5)$ | $G_{n}(5)$ | $G_{n}(5.5)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 51 | 0.19109 | 0.37144 | 0.53206 | 0.66704 | 0.77405 | 0.85404 | 0.91041 | 0.94783 | 0.97123 | 0.98500 | 0.99261 |
| 52 | 0.18929 | 0.36813 | 0.52777 | 0.66236 | 0.76953 | 0.85009 | 0.90725 | 0.94550 | 0.96963 | 0.98398 | 0.99201 |
| 53 | 0.18754 | 0.36492 | 0.52357 | 0.65777 | 0.76508 | 0.84618 | 0.90410 | 0.94315 | 0.96801 | 0.98293 | 0.99138 |
| 54 | 0.18584 | 0.36178 | 0.51948 | 0.65327 | 0.76070 | 0.84231 | 0.90095 | 0.94079 | 0.96636 | 0.98186 | 0.99073 |
| 55 | 0.18418 | 0.35872 | 0.51548 | 0.64886 | 0.75639 | 0.83848 | 0.89782 | 0.93841 | 0.96468 | 0.98076 | 0.99005 |
| 56 | 0.18256 | 0.35574 | 0.51156 | 0.64454 | 0.75213 | 0.83468 | 0.89469 | 0.93603 | 0.96299 | 0.97963 | 0.98935 |
| 57 | 0.18099 | 0.35284 | 0.50774 | 0.64029 | 0.74795 | 0.83091 | 0.89157 | 0.93363 | 0.96127 | 0.97848 | 0.98863 |
| 58 | 0.17946 | 0.35000 | 0.50400 | 0.63613 | 0.74382 | 0.82719 | 0.88847 | 0.93123 | 0.95954 | 0.97731 | 0.98789 |
| 59 | 0.17797 | 0.34723 | 0.50034 | 0.63204 | 0.73975 | 0.82350 | 0.88538 | 0.92882 | 0.95778 | 0.97612 | 0.98712 |
| 60 | 0.17651 | 0.34452 | 0.49675 | 0.62803 | 0.73575 | 0.81985 | 0.88231 | 0.92641 | 0.95602 | 0.97490 | 0.98634 |
| 61 | 0.17509 | 0.34188 | 0.49325 | 0.62409 | 0.73180 | 0.81623 | 0.87924 | 0.92399 | 0.95423 | 0.97366 | 0.98553 |
| 62 | 0.17370 | 0.33930 | 0.48981 | 0.62022 | 0.72791 | 0.81266 | 0.87620 | 0.92158 | 0.95243 | 0.97241 | 0.98471 |
| 63 | 0.17234 | 0.33677 | 0.48645 | 0.61642 | 0.72407 | 0.80911 | 0.87317 | 0.91916 | 0.95062 | 0.97113 | 0.98386 |
| 64 | 0.17102 | 0.33430 | 0.48315 | 0.61269 | 0.72029 | 0.80561 | 0.87015 | 0.91674 | 0.94880 | 0.96984 | 0.98300 |
| 65 | 0.16972 | 0.33189 | 0.47992 | 0.60902 | 0.71657 | 0.80214 | 0.86716 | 0.91432 | 0.94696 | 0.96853 | 0.98211 |
| 66 | 0.16846 | 0.32952 | 0.47675 | 0.60541 | 0.71289 | 0.79871 | 0.86418 | 0.91190 | 0.94512 | 0.96720 | 0.98121 |
| 67 | 0.16722 | 0.32720 | 0.47364 | 0.60187 | 0.70927 | 0.79531 | 0.86121 | 0.90948 | 0.94326 | 0.96586 | 0.98029 |
| 68 | 0.16601 | 0.32494 | 0.47060 | 0.59839 | 0.70570 | 0.79195 | 0.85827 | 0.90707 | 0.94140 | 0.96450 | 0.97936 |
| 69 | 0.16482 | 0.32272 | 0.46761 | 0.59496 | 0.70218 | 0.78862 | 0.85535 | 0.90465 | 0.93953 | 0.96313 | 0.97841 |
| 70 | 0.16366 | 0.32054 | 0.46467 | 0.59160 | 0.69871 | 0.78533 | 0.85244 | 0.90225 | 0.93765 | 0.96175 | 0.97744 |
| 71 | 0.16253 | 0.31841 | 0.46180 | 0.58828 | 0.69528 | 0.78207 | 0.84955 | 0.89985 | 0.93577 | 0.96035 | 0.97646 |
| 72 | 0.16142 | 0.31632 | 0.45897 | 0.58502 | 0.69190 | 0.77884 | 0.84668 | 0.89745 | 0.93388 | 0.95894 | 0.97546 |
| 73 | 0.16033 | 0.31427 | 0.45619 | 0.58182 | 0.68857 | 0.77565 | 0.84383 | 0.89506 | 0.93198 | 0.95752 | 0.97445 |
| 74 | 0.15926 | 0.31226 | 0.45347 | 0.57866 | 0.68529 | 0.77250 | 0.84100 | 0.89267 | 0.93009 | 0.95609 | 0.97342 |
| 75 | 0.15821 | 0.31028 | 0.45079 | 0.57556 | 0.68204 | 0.76937 | 0.83819 | 0.89030 | 0.92819 | 0.95464 | 0.97239 |
| 76 | 0.15718 | 0.30835 | 0.44816 | 0.57250 | 0.67884 | 0.76628 | 0.83540 | 0.88793 | 0.92628 | 0.95319 | 0.97134 |
| 77 | 0.15618 | 0.30645 | 0.44557 | 0.56949 | 0.67568 | 0.76322 | 0.83263 | 0.88556 | 0.92437 | 0.95173 | 0.97027 |
| 78 | 0.15519 | 0.30458 | 0.44303 | 0.56653 | 0.67257 | 0.76019 | 0.82988 | 0.88321 | 0.92247 | 0.95026 | 0.96920 |
| 79 | 0.15422 | 0.30275 | 0.44053 | 0.56361 | 0.66949 | 0.75720 | 0.82715 | 0.88086 | 0.92056 | 0.94879 | 0.96811 |
| 80 | 0.15327 | 0.30095 | 0.43807 | 0.56073 | 0.66645 | 0.75423 | 0.82444 | 0.87852 | 0.91865 | 0.94730 | 0.96701 |
| 81 | 0.15233 | 0.29919 | 0.43565 | 0.55790 | 0.66346 | 0.75130 | 0.82175 | 0.87620 | 0.91674 | 0.94581 | 0.96590 |
| 82 | 0.15142 | 0.29745 | 0.43328 | 0.55511 | 0.66050 | 0.74839 | 0.81908 | 0.87388 | 0.91482 | 0.94432 | 0.96478 |
| 83 | 0.15052 | 0.29574 | 0.43094 | 0.55236 | 0.65757 | 0.74552 | 0.81643 | 0.87157 | 0.91292 | 0.94281 | 0.96365 |
| 84 | 0.14963 | 0.29406 | 0.42863 | 0.54965 | 0.65469 | 0.74267 | 0.81379 | 0.86927 | 0.91101 | 0.94130 | 0.96252 |
| 85 | 0.14876 | 0.29242 | 0.42637 | 0.54698 | 0.65184 | 0.73986 | 0.81118 | 0.86697 | 0.90910 | 0.93979 | 0.96137 |
| 86 | 0.14791 | 0.29079 | 0.42414 | 0.54435 | 0.64903 | 0.73707 | 0.80859 | 0.86469 | 0.90719 | 0.93827 | 0.96021 |
| 87 | 0.14707 | 0.28920 | 0.42194 | 0.54175 | 0.64625 | 0.73431 | 0.80602 | 0.86242 | 0.90529 | 0.93675 | 0.95905 |
| 88 | 0.14624 | 0.28763 | 0.41978 | 0.53919 | 0.64350 | 0.73158 | 0.80346 | 0.86017 | 0.90339 | 0.93522 | 0.95788 |
| 89 | 0.14543 | 0.28608 | 0.41765 | 0.53667 | 0.64079 | 0.72887 | 0.80093 | 0.85792 | 0.90149 | 0.93369 | 0.95670 |
| 90 | 0.14463 | 0.28456 | 0.41555 | 0.53418 | 0.63811 | 0.72620 | 0.79841 | 0.85568 | 0.89959 | 0.93216 | 0.95551 |
| 91 | 0.14385 | 0.28307 | 0.41349 | 0.53172 | 0.63546 | 0.72355 | 0.79592 | 0.85345 | 0.89770 | 0.93063 | 0.95432 |
| 92 | 0.14307 | 0.28160 | 0.41145 | 0.52930 | 0.63285 | 0.72092 | 0.79344 | 0.85124 | 0.89581 | 0.92909 | 0.95312 |
| 93 | 0.14231 | 0.28015 | 0.40944 | 0.52691 | 0.63026 | 0.71832 | 0.79098 | 0.84903 | 0.89393 | 0.92755 | 0.95192 |
| 94 | 0.14156 | 0.27872 | 0.40747 | 0.52455 | 0.62771 | 0.71575 | 0.78854 | 0.84684 | 0.89205 | 0.92601 | 0.95070 |
| 95 | 0.14083 | 0.27731 | 0.40552 | 0.52222 | 0.62518 | 0.71321 | 0.78612 | 0.84465 | 0.89017 | 0.92446 | 0.94949 |
| 96 | 0.14010 | 0.27593 | 0.40360 | 0.51993 | 0.62269 | 0.71068 | 0.78372 | 0.84248 | 0.88830 | 0.92292 | 0.94827 |
| 97 | 0.13939 | 0.27456 | 0.40170 | 0.51766 | 0.62022 | 0.70819 | 0.78134 | 0.84032 | 0.88643 | 0.92137 | 0.94704 |
| 98 | 0.13868 | 0.27322 | 0.39983 | 0.51542 | 0.61778 | 0.70571 | 0.77897 | 0.83818 | 0.88457 | 0.91983 | 0.94581 |
| 99 | 0.13799 | 0.27190 | 0.39799 | 0.51321 | 0.61537 | 0.70326 | 0.77663 | 0.83604 | 0.88271 | 0.91828 | 0.94457 |
| 100 | 0.13731 | 0.27059 | 0.39617 | 0.51103 | 0.61299 | 0.70084 | 0.77430 | 0.83391 | 0.88086 | 0.91674 | 0.94333 |

## 11 Table of $H_{n}(d)$

$H_{n}(d)$ : the probability that the signed visible rounding difference between the sum of $n$ rounded terms and the rounded sum is equal to $d$ times the rounding precision. $H_{n}(d)=F_{n}\left(d+\frac{1}{2}\right)-F_{n}\left(d-\frac{1}{2}\right)$, so by symmetry $H_{n}(-d)=H_{n}(d)$. If instead percentages are used making the sum $100 \%$, then the probability is about $J_{n}(d)=H_{n-1}(d)$ times the rounding precision, so using the preceding row.

| $n$ | $H_{n}(0)$ | $H_{n}(1)$ | $H_{n}(2)$ | $H_{n}(3)$ | $H_{n}(4)$ | $H_{n}(5)$ | $H_{n}(6)$ | $H_{n}(7)$ | $H_{n}(8)$ | $H_{n}(9)$ | $H_{n}(10)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 1 | 1 | 0 |  |  |  |  |  |  |  |  |  |
| 2 | 0.75 | 0.125 |  |  |  |  |  |  |  |  |  |
| 3 | 0.66667 | 0.16667 | 0 |  |  |  |  |  |  |  |  |
| 4 | 0.59896 | 0.19792 | 0.00260 |  |  |  |  |  |  |  |  |
| 5 | 0.55 | 0.21667 | 0.00833 | 0 |  |  |  |  |  |  |  |
| 6 | 0.51102 | 0.22880 | 0.01567 | 0.00002 |  |  |  |  |  |  |  |
| 7 | 0.47937 | 0.23631 | 0.02381 | 0.00020 | 0 |  |  |  |  |  |  |
| 8 | 0.45292 | 0.24078 | 0.03213 | 0.00063 | 0.00000 |  |  |  |  |  |  |
| 9 | 0.43042 | 0.24315 | 0.04026 | 0.00138 | 0.00000 | 0 |  |  |  |  |  |
| 10 | 0.41096 | 0.24407 | 0.04798 | 0.00245 | 0.00002 | 0.00000 |  |  |  |  |  |
| 11 | 0.39393 | 0.24396 | 0.05520 | 0.00382 | 0.00005 | 0.00000 | 0 |  |  |  |  |
| 12 | 0.37884 | 0.24313 | 0.06187 | 0.00546 | 0.00012 | 0.00000 | 0.00000 |  |  |  |  |
| 13 | 0.36537 | 0.24179 | 0.06797 | 0.00731 | 0.00024 | 0.00000 | 0.00000 | 0 |  |  |  |
| 14 | 0.35324 | 0.24008 | 0.07354 | 0.00934 | 0.00041 | 0.00000 | 0.00000 | 0.00000 |  |  |  |
| 15 | 0.34224 | 0.23812 | 0.07860 | 0.01150 | 0.00065 | 0.00001 | 0.00000 | 0.00000 | 0 |  |  |
| 16 | 0.33221 | 0.23599 | 0.08317 | 0.01376 | 0.00095 | 0.00002 | 0.00000 | 0.00000 | 0.00000 |  |  |
| 17 | 0.32301 | 0.23374 | 0.08731 | 0.01607 | 0.00133 | 0.00004 | 0.00000 | 0.00000 | 0.00000 | 0 |  |
| 18 | 0.31453 | 0.23141 | 0.09105 | 0.01842 | 0.00178 | 0.00007 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |  |
| 19 | 0.30669 | 0.22905 | 0.09442 | 0.02078 | 0.00229 | 0.00011 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0 |
| 20 | 0.29941 | 0.22666 | 0.09746 | 0.02313 | 0.00287 | 0.00017 | 0.00000 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 21 | 0.29262 | 0.22428 | 0.10019 | 0.02545 | 0.00351 | 0.00024 | 0.00001 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 22 | 0.28628 | 0.22191 | 0.10266 | 0.02774 | 0.00421 | 0.00034 | 0.00001 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 23 | 0.28033 | 0.21957 | 0.10487 | 0.02997 | 0.00495 | 0.00045 | 0.00002 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 24 | 0.27473 | 0.21726 | 0.10687 | 0.03216 | 0.00574 | 0.00058 | 0.00003 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 25 | 0.26946 | 0.21498 | 0.10866 | 0.03428 | 0.00657 | 0.00073 | 0.00004 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 26 | 0.26448 | 0.21275 | 0.11026 | 0.03634 | 0.00744 | 0.00091 | 0.00006 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 27 | 0.25977 | 0.21055 | 0.11170 | 0.03833 | 0.00833 | 0.00111 | 0.00009 | 0.00000 | 0.00000 | 0.00000 | 0.00000 |
| 28 | 0.25530 | 0.20841 | 0.11299 | 0.04025 | 0.00925 | 0.00133 | 0.00012 | 0.00001 | 0.00000 | 0.00000 | 0.00000 |
| 29 | 0.25105 | 0.20631 | 0.11414 | 0.04211 | 0.01019 | 0.00157 | 0.00015 | 0.00001 | 0.00000 | 0.00000 | 0.00000 |
| 30 | 0.24701 | 0.20425 | 0.11517 | 0.04390 | 0.01114 | 0.00184 | 0.00019 | 0.00001 | 0.00000 | 0.00000 | 0.00000 |
| 31 | 0.24315 | 0.20224 | 0.11608 | 0.04561 | 0.01210 | 0.00212 | 0.00024 | 0.00002 | 0.00000 | 0.00000 | 0.00000 |
| 32 | 0.23948 | 0.20028 | 0.11689 | 0.04727 | 0.01308 | 0.00243 | 0.00030 | 0.00002 | 0.00000 | 0.00000 | 0.00000 |
| 33 | 0.23596 | 0.19837 | 0.11760 | 0.04885 | 0.01406 | 0.00275 | 0.00036 | 0.00003 | 0.00000 | 0.00000 | 0.00000 |
| 34 | 0.23259 | 0.19649 | 0.11823 | 0.05037 | 0.01504 | 0.00310 | 0.00043 | 0.00004 | 0.00000 | 0.00000 | 0.00000 |
| 35 | 0.22937 | 0.19466 | 0.11879 | 0.05183 | 0.01602 | 0.00346 | 0.00051 | 0.00005 | 0.00000 | 0.00000 | 0.00000 |
| 36 | 0.22627 | 0.19288 | 0.11927 | 0.05323 | 0.01699 | 0.00383 | 0.00060 | 0.00006 | 0.00000 | 0.00000 | 0.00000 |
| 37 | 0.22330 | 0.19113 | 0.11968 | 0.05456 | 0.01797 | 0.00422 | 0.00070 | 0.00008 | 0.00001 | 0.00000 | 0.00000 |
| 38 | 0.22044 | 0.18943 | 0.12003 | 0.05585 | 0.01893 | 0.00463 | 0.00080 | 0.00010 | 0.00001 | 0.00000 | 0.00000 |
| 39 | 0.21769 | 0.18777 | 0.12033 | 0.05707 | 0.01989 | 0.00504 | 0.00092 | 0.00012 | 0.00001 | 0.00000 | 0.00000 |
| 40 | 0.21504 | 0.18614 | 0.12058 | 0.05824 | 0.02084 | 0.00547 | 0.00104 | 0.00014 | 0.00001 | 0.00000 | 0.00000 |
| 41 | 0.21248 | 0.18455 | 0.12079 | 0.05937 | 0.02178 | 0.00591 | 0.00117 | 0.00017 | 0.00002 | 0.00000 | 0.00000 |
| 42 | 0.21001 | 0.18300 | 0.12095 | 0.06044 | 0.02271 | 0.00636 | 0.00132 | 0.00020 | 0.00002 | 0.00000 | 0.00000 |
| 43 | 0.20763 | 0.18148 | 0.12107 | 0.06146 | 0.02362 | 0.00682 | 0.00147 | 0.00023 | 0.00003 | 0.00000 | 0.00000 |
| 44 | 0.20533 | 0.18000 | 0.12116 | 0.06244 | 0.02452 | 0.00729 | 0.00162 | 0.00027 | 0.00003 | 0.00000 | 0.00000 |
| 45 | 0.20310 | 0.17855 | 0.12121 | 0.06338 | 0.02541 | 0.00776 | 0.00179 | 0.00031 | 0.00004 | 0.00000 | 0.00000 |
| 46 | 0.20094 | 0.17713 | 0.12124 | 0.06427 | 0.02628 | 0.00824 | 0.00196 | 0.00035 | 0.00005 | 0.00000 | 0.00000 |
| 47 | 0.19885 | 0.17574 | 0.12124 | 0.06513 | 0.02714 | 0.00872 | 0.00215 | 0.00040 | 0.00006 | 0.00001 | 0.00000 |
| 48 | 0.19682 | 0.17439 | 0.12121 | 0.06595 | 0.02798 | 0.00921 | 0.00233 | 0.00045 | 0.00007 | 0.00001 | 0.00000 |
| 49 | 0.19485 | 0.17306 | 0.12116 | 0.06673 | 0.02881 | 0.00970 | 0.00253 | 0.00051 | 0.00008 | 0.00001 | 0.00000 |
| 50 | 0.19295 | 0.17176 | 0.12108 | 0.06747 | 0.02962 | 0.01019 | 0.00273 | 0.00057 | 0.00009 | 0.00001 | 0.00000 |

$H_{n}(d)$ continued: the probability that the signed visible rounding difference between the sum of $n$ rounded terms and the rounded sum is equal to $d$ times the rounding precision. $H_{n}(-d)=H_{n}(d)$. For percentages, the probability is about $J_{n}(d)=H_{n-1}(d)$ times the rounding precision.

| $n$ | $H_{n}(0)$ | $H_{n}(1)$ | $H_{n}(2)$ | $H_{n}(3)$ | $H_{n}(4)$ | $H_{n}(5)$ | $H_{n}(6)$ | $H_{n}(7)$ | $H_{n}(8)$ | $H_{n}(9)$ | $H_{n}(10)$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 51 | 0.19109 | 0.17049 | 0.12099 | 0.06818 | 0.03041 | 0.01069 | 0.00294 | 0.00063 | 0.00010 | 0.00001 | 0.00000 |
| 52 | 0.18929 | 0.16924 | 0.12088 | 0.06886 | 0.03119 | 0.01119 | 0.00316 | 0.00070 | 0.00012 | 0.00002 | 0.00000 |
| 53 | 0.18754 | 0.16802 | 0.12075 | 0.06951 | 0.03196 | 0.01169 | 0.00338 | 0.00077 | 0.00014 | 0.00002 | 0.00000 |
| 54 | 0.18584 | 0.16682 | 0.12061 | 0.07013 | 0.03270 | 0.01219 | 0.00361 | 0.00085 | 0.00016 | 0.00002 | 0.00000 |
| 55 | 0.18418 | 0.16565 | 0.12045 | 0.07071 | 0.03343 | 0.01269 | 0.00384 | 0.00092 | 0.00018 | 0.00003 | 0.00000 |
| 56 | 0.18256 | 0.16450 | 0.12028 | 0.07128 | 0.03415 | 0.01318 | 0.00408 | 0.00101 | 0.00020 | 0.00003 | 0.00000 |
| 57 | 0.18099 | 0.16337 | 0.12010 | 0.07181 | 0.03485 | 0.01368 | 0.00433 | 0.00110 | 0.00022 | 0.00004 | 0.00000 |
| 58 | 0.17946 | 0.16227 | 0.11991 | 0.07233 | 0.03553 | 0.01418 | 0.00457 | 0.00119 | 0.00025 | 0.00004 | 0.00001 |
| 59 | 0.17797 | 0.16119 | 0.11971 | 0.07281 | 0.03620 | 0.01467 | 0.00483 | 0.00128 | 0.00027 | 0.00005 | 0.00001 |
| 60 | 0.17651 | 0.16012 | 0.11950 | 0.07328 | 0.03686 | 0.01516 | 0.00508 | 0.00138 | 0.00030 | 0.00005 | 0.00001 |
| 61 | 0.17509 | 0.15908 | 0.11928 | 0.07372 | 0.03749 | 0.01565 | 0.00534 | 0.00149 | 0.00033 | 0.00006 | 0.00001 |
| 62 | 0.17370 | 0.15806 | 0.11905 | 0.07414 | 0.03812 | 0.01614 | 0.00561 | 0.00159 | 0.00037 | 0.00007 | 0.00001 |
| 63 | 0.17234 | 0.15705 | 0.11881 | 0.07455 | 0.03873 | 0.01662 | 0.00587 | 0.00170 | 0.00040 | 0.00008 | 0.00001 |
| 64 | 0.17102 | 0.15607 | 0.11857 | 0.07493 | 0.03932 | 0.01710 | 0.00614 | 0.00182 | 0.00044 | 0.00009 | 0.00001 |
| 65 | 0.16972 | 0.15510 | 0.11832 | 0.07529 | 0.03990 | 0.01758 | 0.00642 | 0.00193 | 0.00048 | 0.00010 | 0.00002 |
| 66 | 0.16846 | 0.15415 | 0.11807 | 0.07564 | 0.04047 | 0.01805 | 0.00669 | 0.00205 | 0.00052 | 0.00011 | 0.00002 |
| 67 | 0.16722 | 0.15321 | 0.11781 | 0.07597 | 0.04102 | 0.01852 | 0.00697 | 0.00218 | 0.00056 | 0.00012 | 0.00002 |
| 68 | 0.16601 | 0.15229 | 0.11755 | 0.07628 | 0.04156 | 0.01898 | 0.00725 | 0.00231 | 0.00061 | 0.00013 | 0.00002 |
| 69 | 0.16482 | 0.15139 | 0.11729 | 0.07658 | 0.04209 | 0.01944 | 0.00753 | 0.00244 | 0.00066 | 0.00015 | 0.00003 |
| 70 | 0.16366 | 0.15051 | 0.11702 | 0.07687 | 0.04261 | 0.01989 | 0.00781 | 0.00257 | 0.00071 | 0.00016 | 0.00003 |
| 71 | 0.16253 | 0.14963 | 0.11674 | 0.07713 | 0.04311 | 0.02035 | 0.00809 | 0.00270 | 0.00076 | 0.00018 | 0.00003 |
| 72 | 0.16142 | 0.14878 | 0.11647 | 0.07739 | 0.04360 | 0.02079 | 0.00838 | 0.00284 | 0.00081 | 0.00019 | 0.00004 |
| 73 | 0.16033 | 0.14793 | 0.11619 | 0.07763 | 0.04408 | 0.02123 | 0.00866 | 0.00298 | 0.00087 | 0.00021 | 0.00004 |
| 74 | 0.15926 | 0.14710 | 0.11591 | 0.07786 | 0.04454 | 0.02167 | 0.00895 | 0.00313 | 0.00092 | 0.00023 | 0.00005 |
| 75 | 0.15821 | 0.14629 | 0.11563 | 0.07808 | 0.04500 | 0.02210 | 0.00923 | 0.00328 | 0.00098 | 0.00025 | 0.00005 |
| 76 | 0.15718 | 0.14549 | 0.11534 | 0.07828 | 0.04544 | 0.02253 | 0.00952 | 0.00342 | 0.00104 | 0.00027 | 0.00006 |
| 77 | 0.15618 | 0.14470 | 0.11506 | 0.07848 | 0.04587 | 0.02295 | 0.00981 | 0.00358 | 0.00111 | 0.00029 | 0.00006 |
| 78 | 0.15519 | 0.14392 | 0.11477 | 0.07866 | 0.04629 | 0.02336 | 0.01010 | 0.00373 | 0.00117 | 0.00031 | 0.00007 |
| 79 | 0.15422 | 0.14315 | 0.11448 | 0.07883 | 0.04670 | 0.02378 | 0.01039 | 0.00388 | 0.00124 | 0.00034 | 0.00008 |
| 80 | 0.15327 | 0.14240 | 0.11419 | 0.07899 | 0.04710 | 0.02418 | 0.01067 | 0.00404 | 0.00131 | 0.00036 | 0.00009 |
| 81 | 0.15233 | 0.14166 | 0.11390 | 0.07915 | 0.04749 | 0.02458 | 0.01096 | 0.00420 | 0.00138 | 0.00039 | 0.00009 |
| 82 | 0.15142 | 0.14093 | 0.11361 | 0.07929 | 0.04787 | 0.02498 | 0.01125 | 0.00436 | 0.00145 | 0.00042 | 0.00010 |
| 83 | 0.15052 | 0.14021 | 0.11332 | 0.07943 | 0.04824 | 0.02537 | 0.01153 | 0.00453 | 0.00153 | 0.00044 | 0.00011 |
| 84 | 0.14963 | 0.13950 | 0.11303 | 0.07955 | 0.04861 | 0.02575 | 0.01182 | 0.00469 | 0.00161 | 0.00047 | 0.00012 |
| 85 | 0.14876 | 0.13880 | 0.11274 | 0.07967 | 0.04896 | 0.02614 | 0.01210 | 0.00486 | 0.00168 | 0.00050 | 0.00013 |
| 86 | 0.14791 | 0.13812 | 0.11244 | 0.07978 | 0.04930 | 0.02651 | 0.01239 | 0.00502 | 0.00176 | 0.00054 | 0.00014 |
| 87 | 0.14707 | 0.13744 | 0.11215 | 0.07989 | 0.04964 | 0.02688 | 0.01267 | 0.00519 | 0.00185 | 0.00057 | 0.00015 |
| 88 | 0.14624 | 0.13677 | 0.11186 | 0.07998 | 0.04996 | 0.02725 | 0.01295 | 0.00536 | 0.00193 | 0.00060 | 0.00016 |
| 89 | 0.14543 | 0.13611 | 0.11157 | 0.08007 | 0.05028 | 0.02760 | 0.01324 | 0.00553 | 0.00202 | 0.00064 | 0.00017 |
| 90 | 0.14463 | 0.13546 | 0.11128 | 0.08015 | 0.05059 | 0.02796 | 0.01352 | 0.00571 | 0.00210 | 0.00067 | 0.00019 |
| 91 | 0.14385 | 0.13482 | 0.11099 | 0.08023 | 0.05089 | 0.02831 | 0.01379 | 0.00588 | 0.00219 | 0.00071 | 0.00020 |
| 92 | 0.14307 | 0.13419 | 0.11070 | 0.08030 | 0.05119 | 0.02865 | 0.01407 | 0.00606 | 0.00228 | 0.00075 | 0.00021 |
| 93 | 0.14231 | 0.13357 | 0.11041 | 0.08036 | 0.05147 | 0.02899 | 0.01435 | 0.00623 | 0.00237 | 0.00079 | 0.00023 |
| 94 | 0.14156 | 0.13295 | 0.11012 | 0.08042 | 0.05175 | 0.02933 | 0.01462 | 0.00641 | 0.00246 | 0.00083 | 0.00024 |
| 95 | 0.14083 | 0.13235 | 0.10983 | 0.08047 | 0.05202 | 0.02966 | 0.01490 | 0.00658 | 0.00256 | 0.00087 | 0.00026 |
| 96 | 0.14010 | 0.13175 | 0.10955 | 0.08052 | 0.05229 | 0.02998 | 0.01517 | 0.00676 | 0.00265 | 0.00091 | 0.00028 |
| 97 | 0.13939 | 0.13116 | 0.10926 | 0.08056 | 0.05255 | 0.03030 | 0.01544 | 0.00694 | 0.00275 | 0.00096 | 0.00029 |
| 98 | 0.13868 | 0.13057 | 0.10897 | 0.08060 | 0.05280 | 0.03062 | 0.01571 | 0.00712 | 0.00285 | 0.00100 | 0.00031 |
| 99 | 0.13799 | 0.13000 | 0.10869 | 0.08063 | 0.05304 | 0.03093 | 0.01597 | 0.00730 | 0.00295 | 0.00105 | 0.00033 |
| 100 | 0.13731 | 0.12943 | 0.10841 | 0.08065 | 0.05328 | 0.03124 | 0.01624 | 0.00748 | 0.00305 | 0.00110 | 0.00035 |

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